

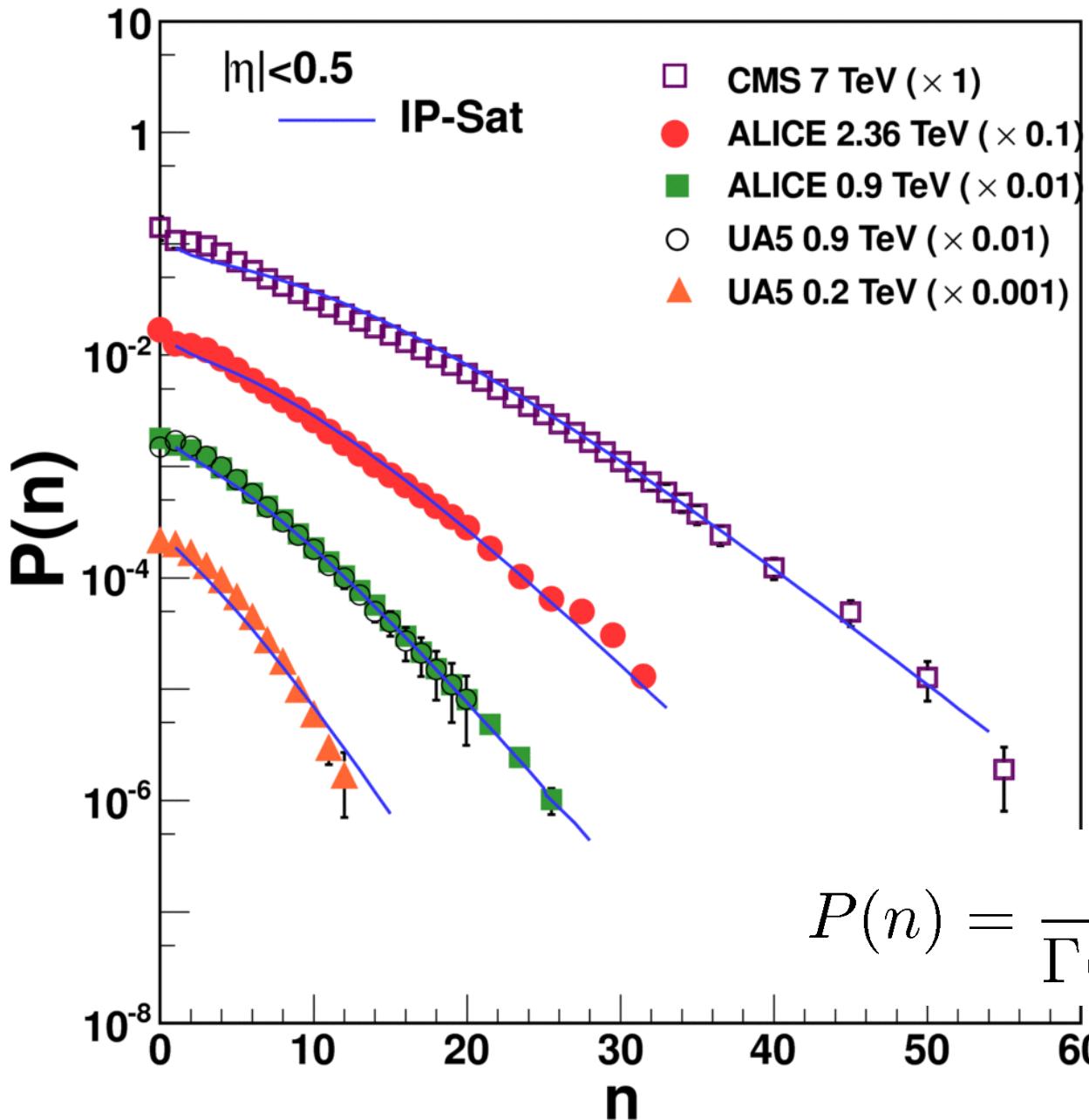
Fluctuations in the CGC: KNO scaling and higher- order eccentricities

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CUNY Graduate Center and
Baruch College/CUNY

RHIC & AGS users meeting, June 12th, 2012

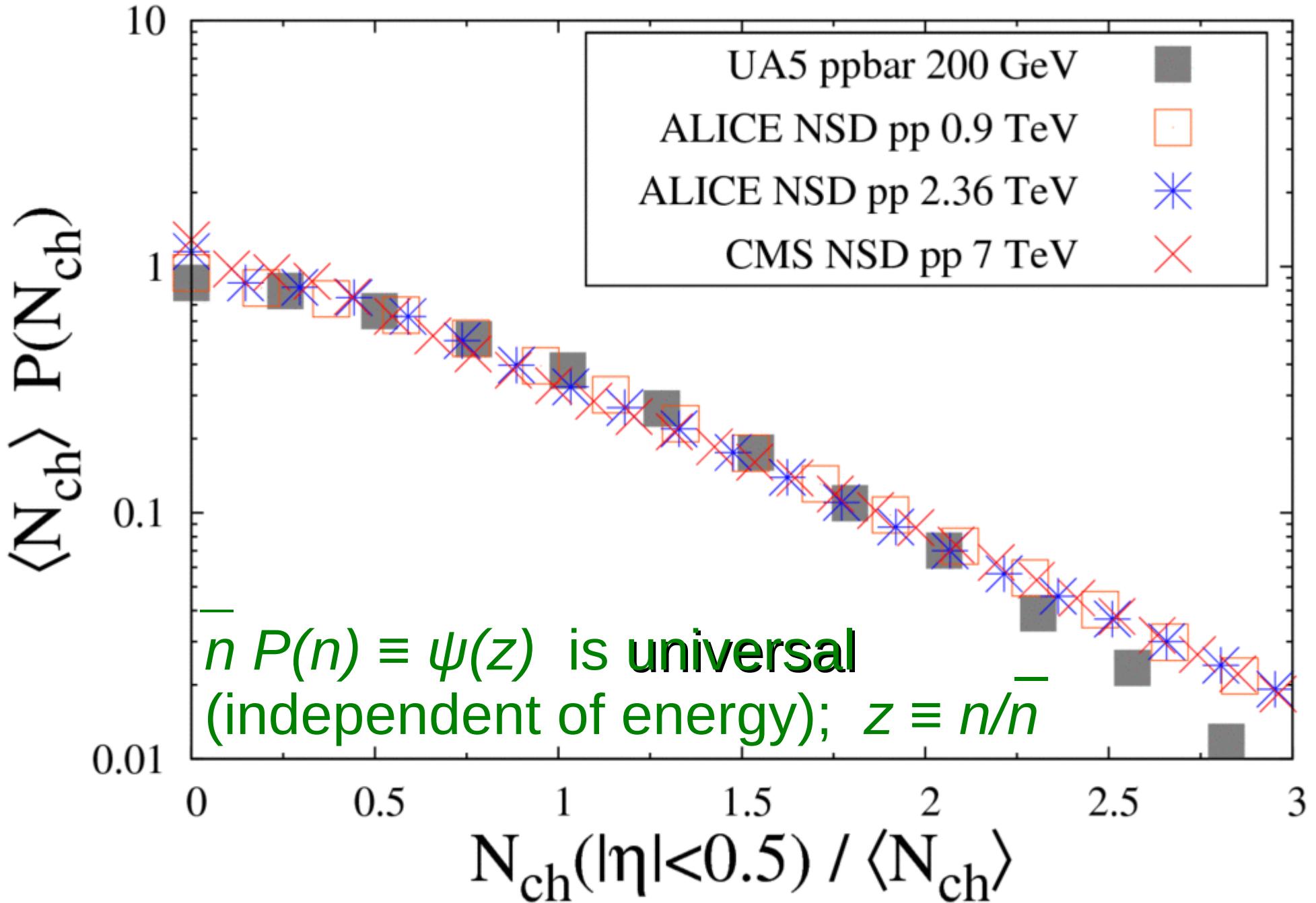
Multiplicity distributions in pp collisions



$P(n)$: negative binomial distribution

$$P(n) = \frac{\Gamma(k+n)}{\Gamma(k)\Gamma(n+1)} \frac{\langle n \rangle^n k^k}{(\langle n \rangle + k)^{n+k}}$$

KNO scaling in high-energy pp data

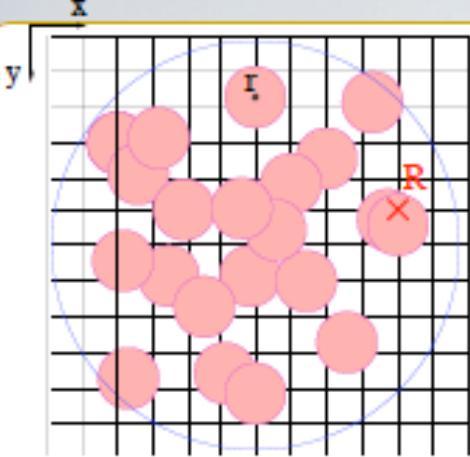


Glauber / geometry fluctuations

(pA and AA collisions)

- fluctuations of transverse positions of valence charges in the colliding nuclei
- fluctuations due to finite probability of inelastic nucleon-nucleon interaction $P(b)$

fluctuations of valence partons in \perp plane



1. Initial conditions for the evolution ($x=0.01$)

$$N(\mathbf{R}) = \sum_{i=1}^A \Theta\left(\sqrt{\frac{\sigma_0}{\pi}} - |\mathbf{R} - \mathbf{r}_i|\right) \rightarrow Q_{s0}^2(\mathbf{R}) = N(\mathbf{R}) Q_{s0, \text{nucl}}^2$$

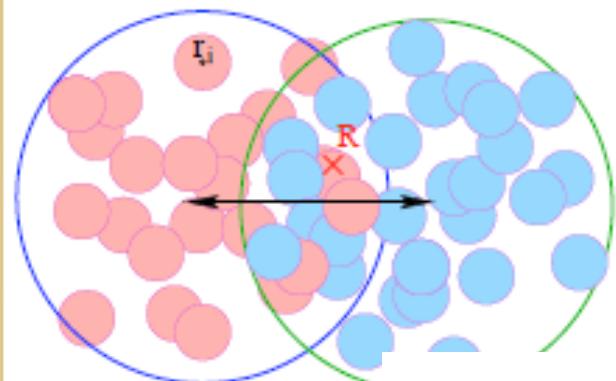
$$\varphi(x_0 = 0.01, k_t, \mathbf{R})$$

rcBK equation
or KLN model

$$\varphi(x, k_t, \mathbf{R})$$

2. Solve local running coupling BK evolution at each transverse point

3 Calculate gluon production at each transverse point according to kt-factorization



INPUT: $\varphi(x = 0.01, k_t)$ FOR A SINGLE NUCLEON:

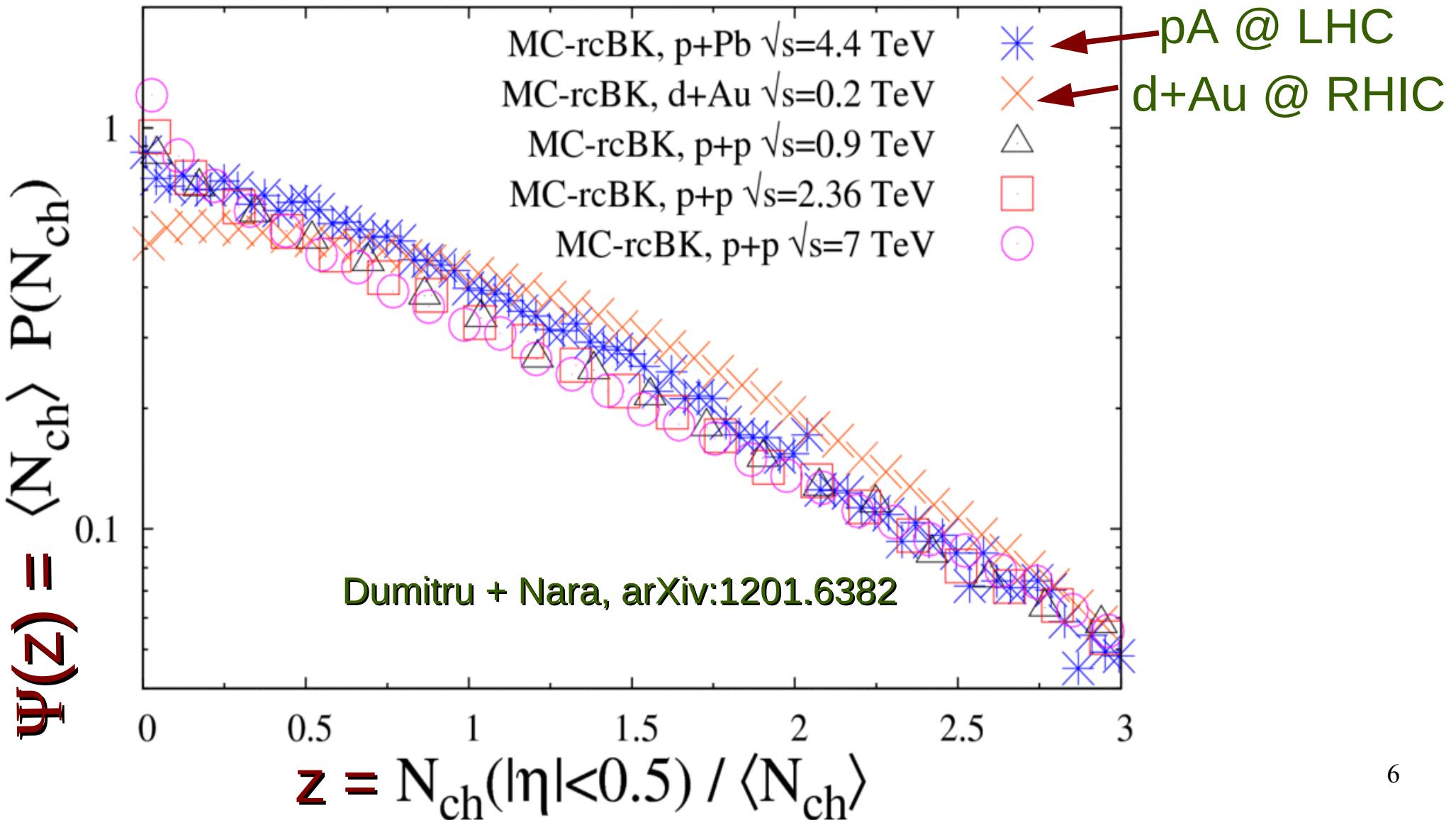
$$N_{\text{part}, A}(\vec{b}) = \sum_{i=1 \dots A} \Theta\left(P(\vec{b} - \vec{r}_i) - \nu_i\right).$$

$$P(b) = 1 - \exp[-\sigma_g T_{pp}(b)], \quad T_{pp}(b) = \int d^2 s T_p(s) T_p(s - b)$$

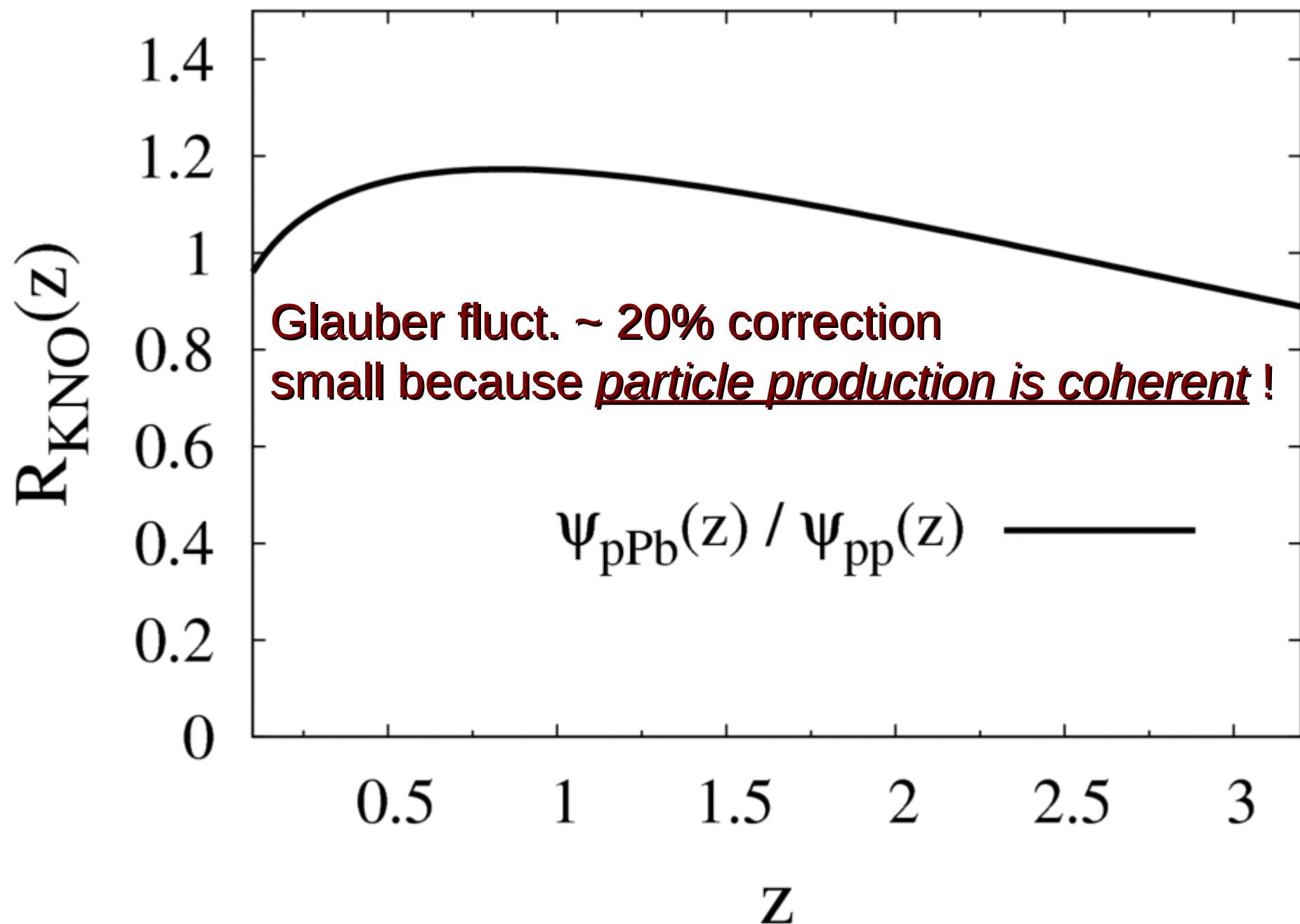
$$T_p(r) = \frac{1}{2\pi B} \exp[-r^2/(2B)] \quad \sigma_{NN}(\sqrt{s}) = \int d^2 b (1 - \exp[-\sigma_g T_{pp}(b)])$$

KNO scaling (even p+Pb approx.; prediction)

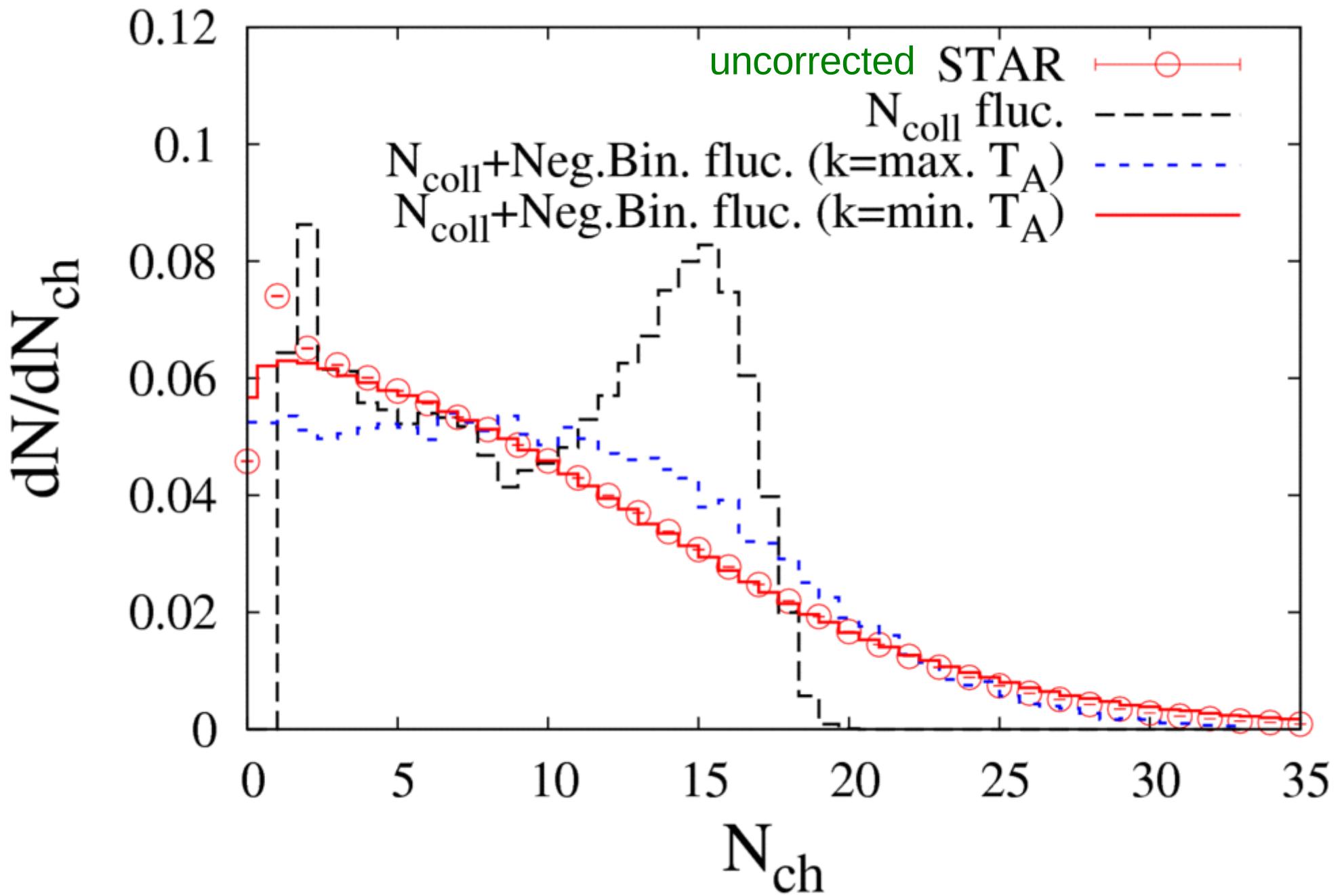
for A+B : $k_{AB} \sim k_{pp} \min(T_A, T_B)$



KNO scaling in p+Pb @ 4.4TeV on linear scale

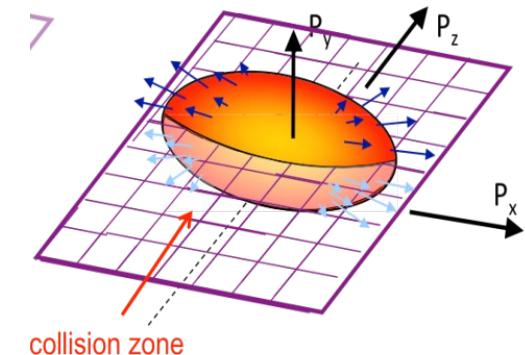
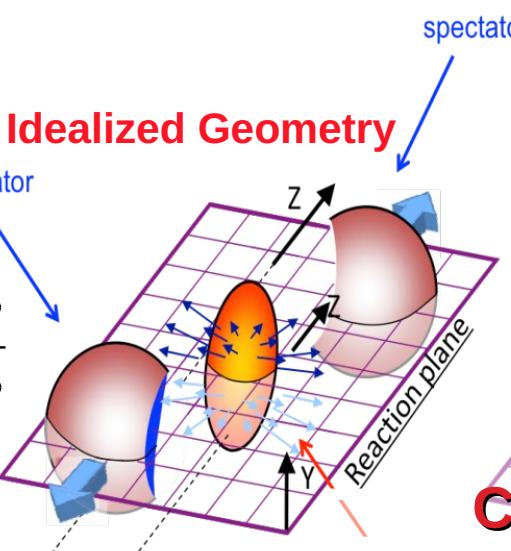


need KNO flucs also for d+Au@RHIC



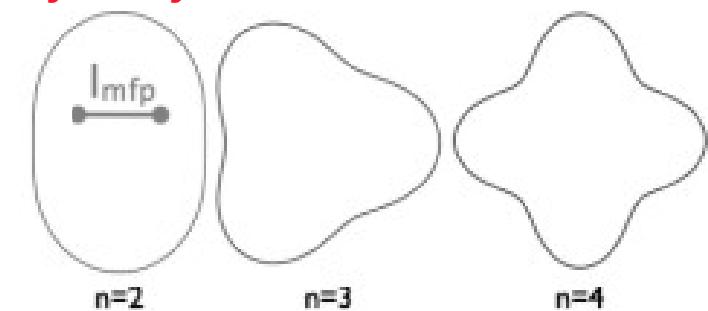
The Flow Probe

$$\epsilon = \frac{\langle y^2 - x^2 \rangle}{\langle y^2 + x^2 \rangle}$$



Control parameters

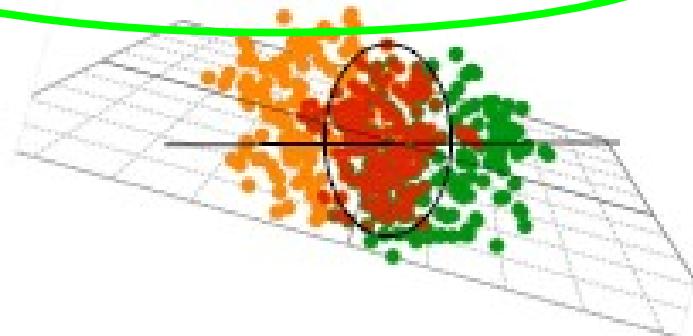
Initial Geometry characterized by many harmonics

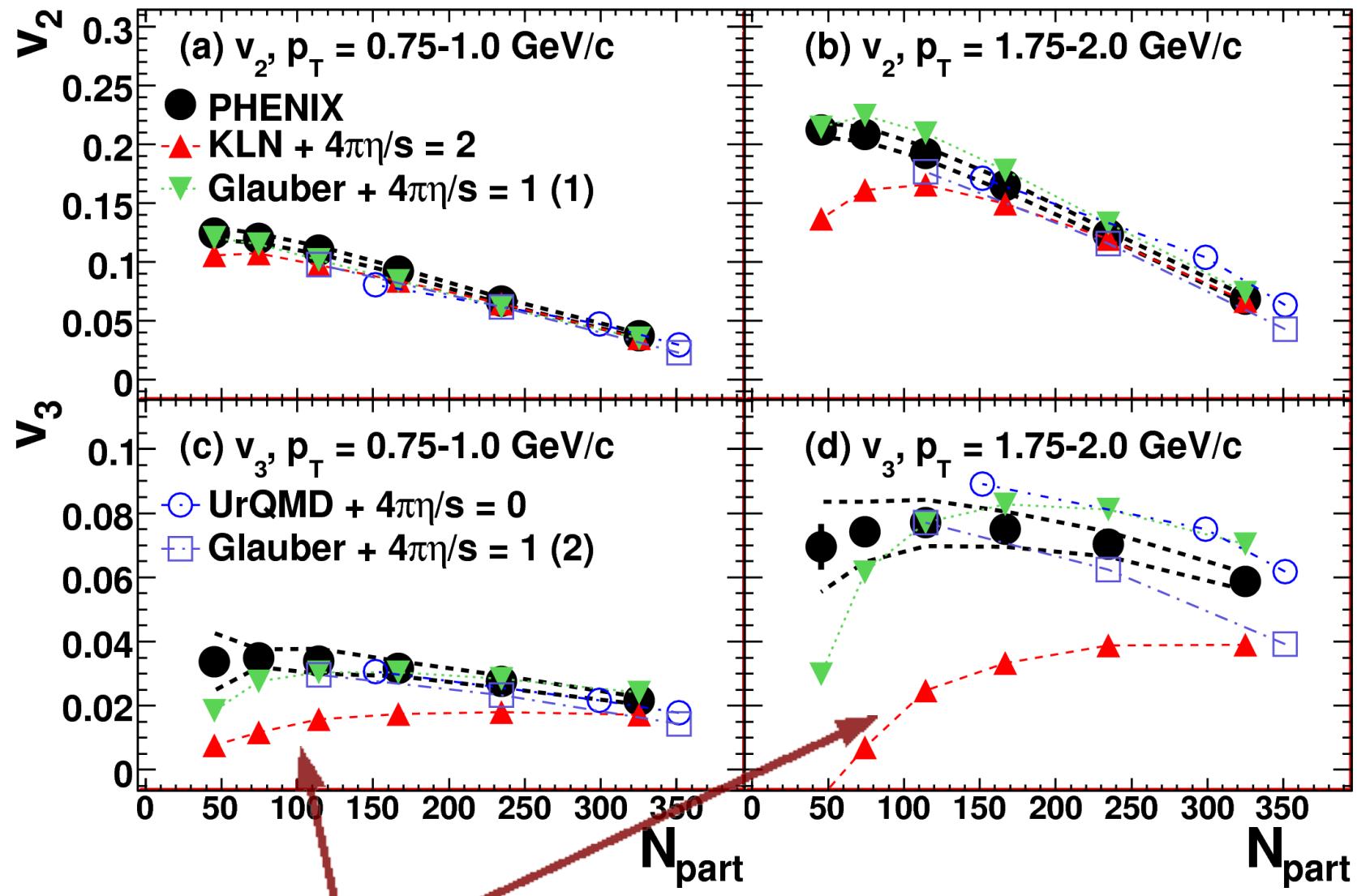


$$\epsilon_n = \frac{\left\langle r^n \cos(n\varphi_{part}) \right\rangle^2 + \left\langle r^n \sin(n\varphi_{part}) \right\rangle^2}{\left\langle r^n \right\rangle^2}$$

Initial eccentricity (and its attendant fluctuations) ϵ_n drive momentum anisotropy v_\parallel with specific scaling properties

Actual collision profiles are not smooth, due to fluctuations!

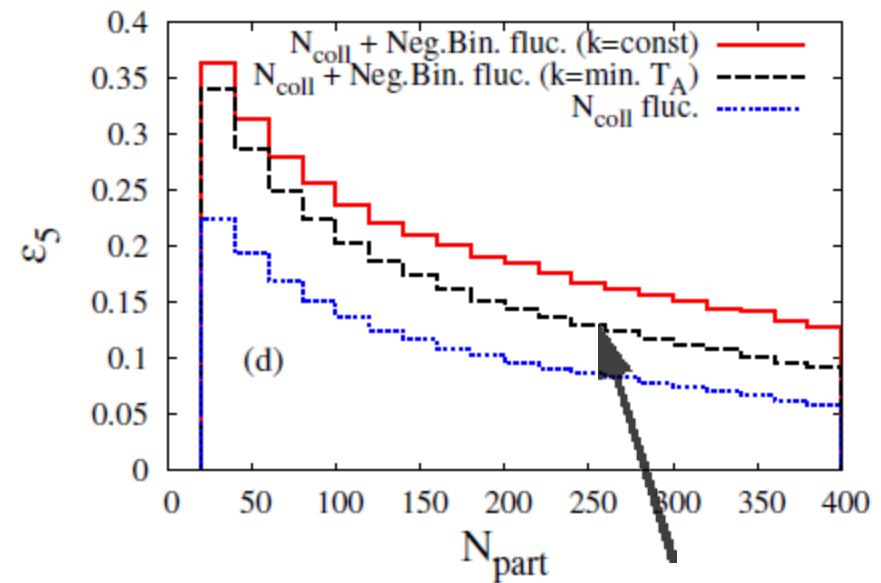
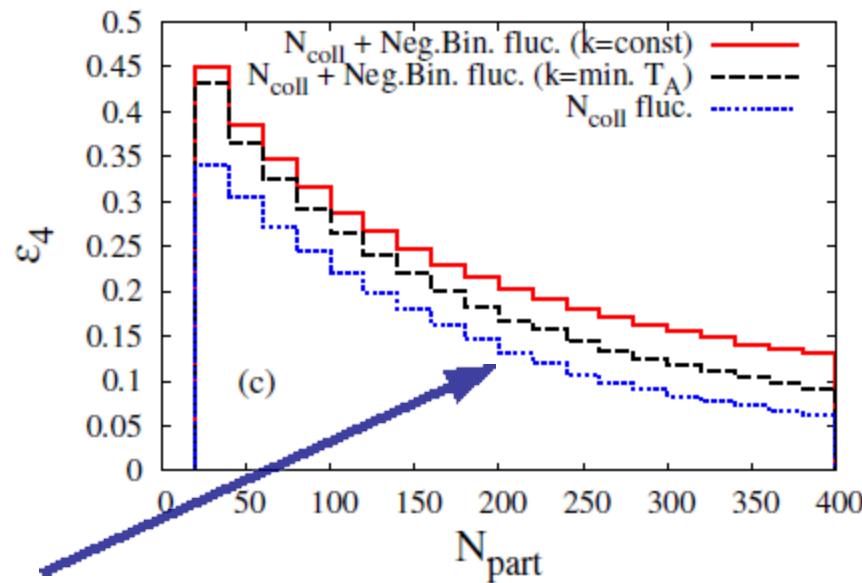
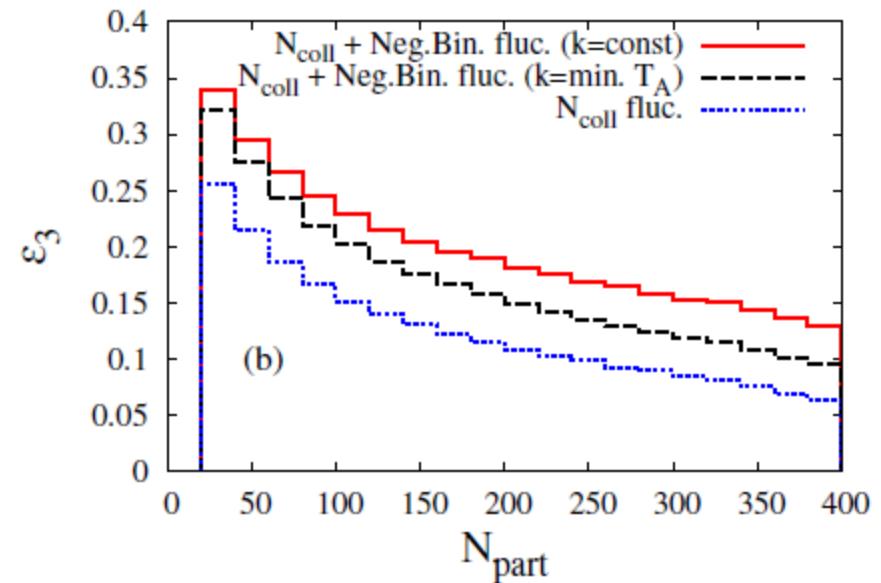
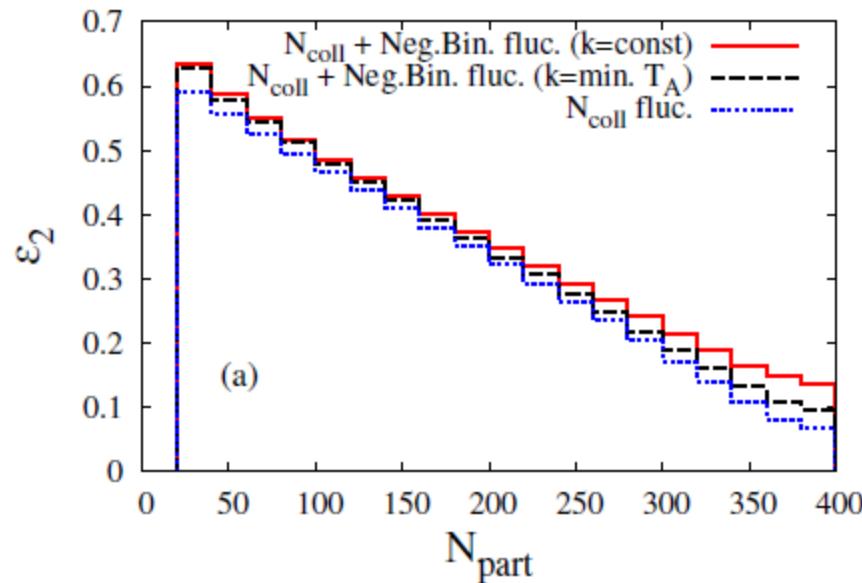




- MC-KLN initial condition (k_T factorization with **Glauber** fluct. only)
leads to underestimate of v_3 !

Eccentricities ε_n in Au+Au

Dumitru + Nara, arXiv:1201.6382



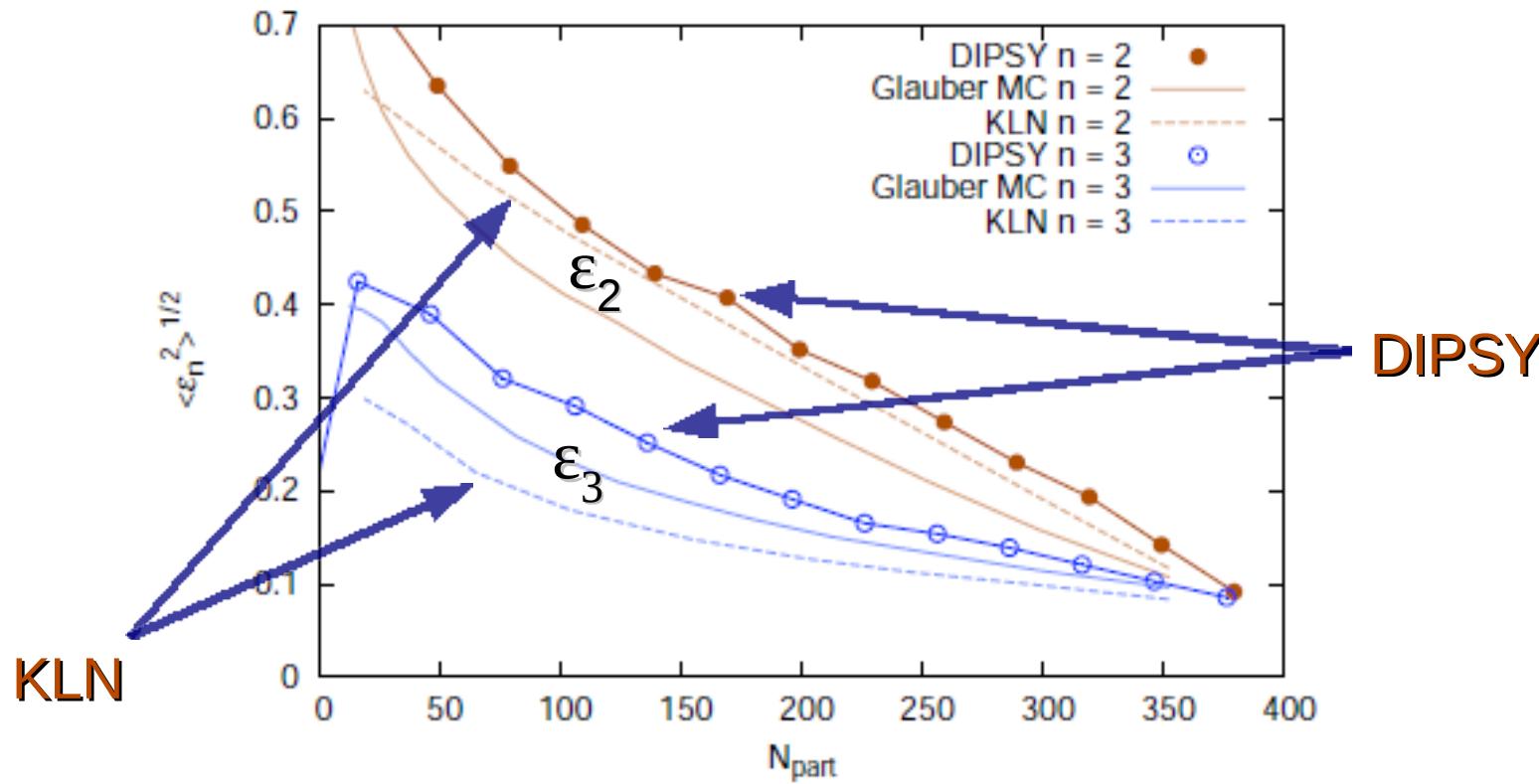
Glauber fluc only

Glauber + NBD
 $k \sim \min(T_A, T_B)^{11}$

DIPSY MC (Lund) w. fluctuations in BFKL ladders

Results: $\varepsilon_2, \varepsilon_3$

C. Flensburg: ISMD 2011,
Hiroshima

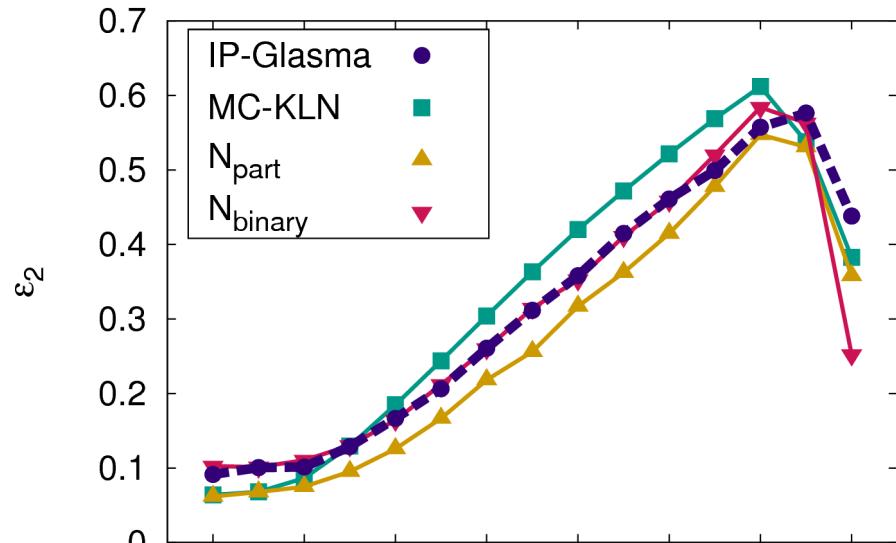


compare DIPSY \longleftrightarrow MC-KLN:
 ε_2 similar, ε_3 larger

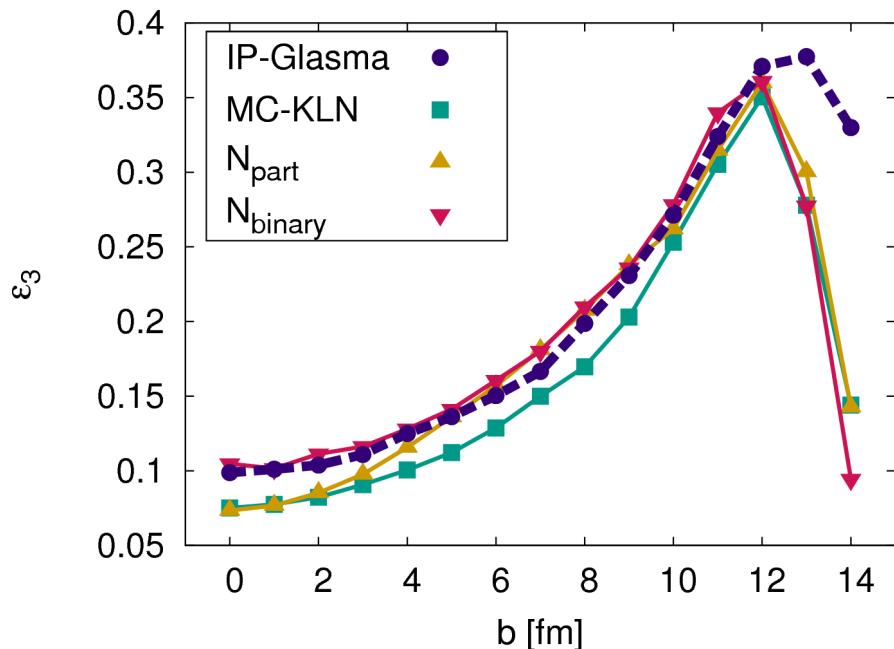
- does DIPSY reproduce KNO in pp, pA ?

Classical YM solutions w/ fluctuating initial conditions (incl. NBD !)

Schenke, Tribedy, Venugopalan,
[arXiv:1202.6646](https://arxiv.org/abs/1202.6646)



- lower ϵ_2 than MC-KLN
- higher ϵ_3



KNO scaling:

Koba, Nielsen, Olesen, NPB 40 (1972) 317

$\bar{n} P(n) \equiv \psi(z)$ is universal (independent of \bar{n} energy); $z \equiv n/\bar{n}$

Note that if $k \ll \bar{n}$, NBD can be written as

$$\bar{n} P(n) dz \sim z^{k-1} e^{-kz} dz , \quad z \equiv n/\bar{n}$$

So, if $k \approx \text{const}$, this leads to KNO scaling !

fit to pp @ LHC: $k / \bar{n} \sim 0.16$ at 2360 GeV

but why is

i) $P(n)$ a NBD ?

ii) $k \ll \bar{n}$?

NBD from MV model

Gelis, Lappi, McLerran:
arXiv:0905.3234

for large nucleus, $A^{1/3} \rightarrow \infty$

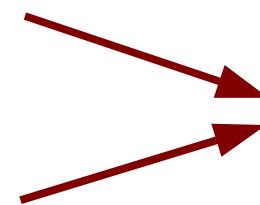
$$S_{\text{MV}} = \int d^2x_\perp \frac{1}{2\mu^2} \rho^a \rho^a + \text{soft YM fields} + \text{coupling of soft} \leftrightarrow \text{hard}$$

$$\begin{aligned} \left\langle \frac{dN}{dy_1 \cdots dy_n} \right\rangle_{\text{conn.}} &= \beta_n \left\langle \frac{dN}{dy_1} \right\rangle \cdots \left\langle \frac{dN}{dy_n} \right\rangle \\ \beta_n &= (n-1)! k^{1-n} \quad \longrightarrow \text{NBD} \\ \bar{n} &= \# \frac{N_c(N_c^2 - 1)}{\alpha_s} Q_s^2 \pi R^2 \\ k &= \# \frac{N_c^2 - 1}{2\pi} Q_s^2 \pi R^2 \sim T_A \end{aligned}$$

So, why KNO then ?

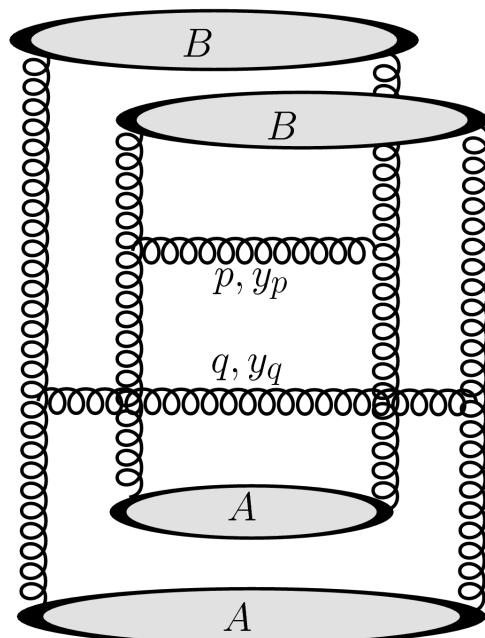
$$\bar{n} = \# \frac{N_c(N_c^2 - 1)}{\alpha_s} Q_s^2 \pi R^2$$

$$k = \# \frac{N_c^2 - 1}{2\pi} Q_s^2 \pi R^2$$

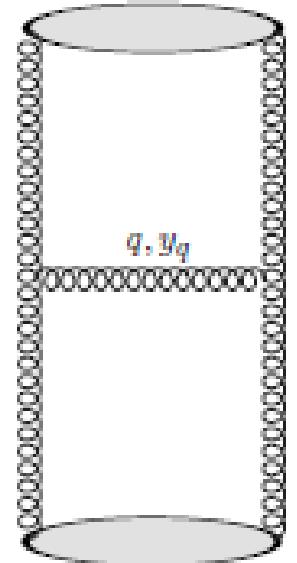
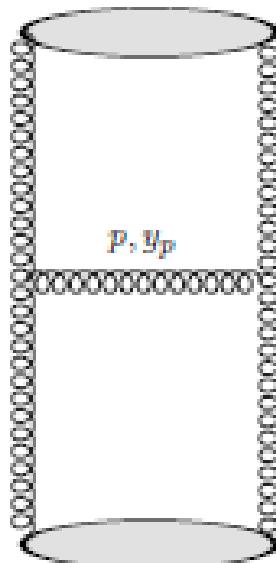


$$\frac{\bar{n}}{k} \sim \frac{N_c}{\alpha_s} \gg 1$$

Why is $k = O(\alpha_s^0)$?



← →
same order
in α_s



KNO scaling emerges if

- i) Gaussian action
- ii) high occupation number

How about

- i) quantum evolution
- ii) corrections to Gaussian action ?

Beyond MV action...

Dumitru, Jalilian-Marian, E.P.
PRD 2011

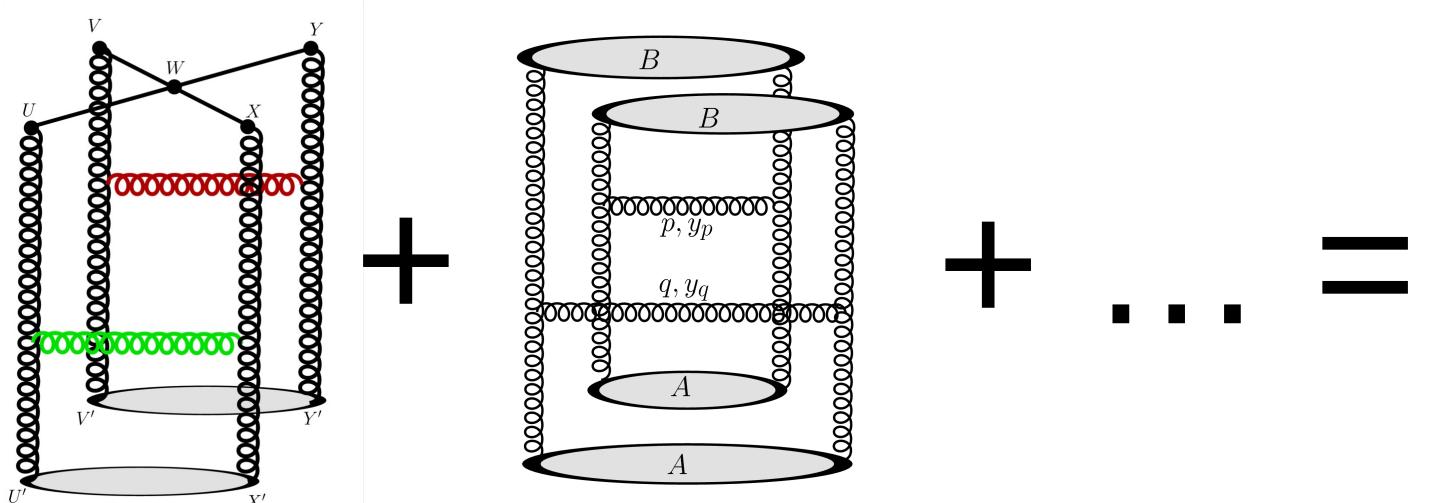
$$S = \int d^2x_\perp \left\{ \frac{1}{2\mu^2} \rho^a \rho^a + \frac{1}{\kappa_4} (\delta^{ab} \delta^{cd} + \delta^{ac} \delta^{bd} + \delta^{ad} \delta^{bc}) \rho^a \rho^b \rho^c \rho^d \right\}$$

+ soft YM fields + coupling of soft \leftrightarrow hard

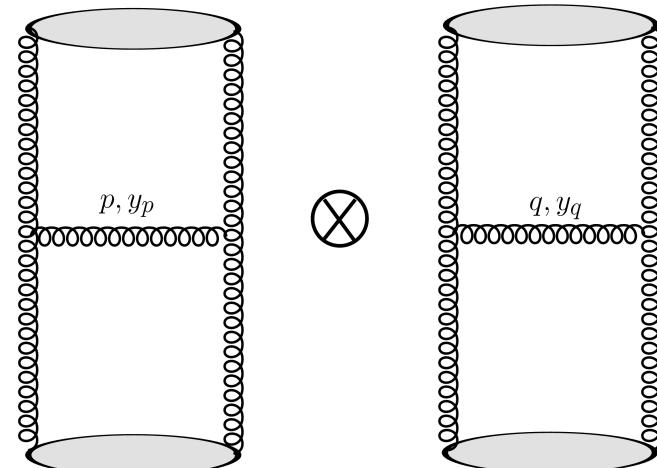
- $\mu^2 \sim g^2 A^{1/3}; \quad \kappa_4 \sim g^4 A$

Recalculate width k^{-1} of mult. distribution

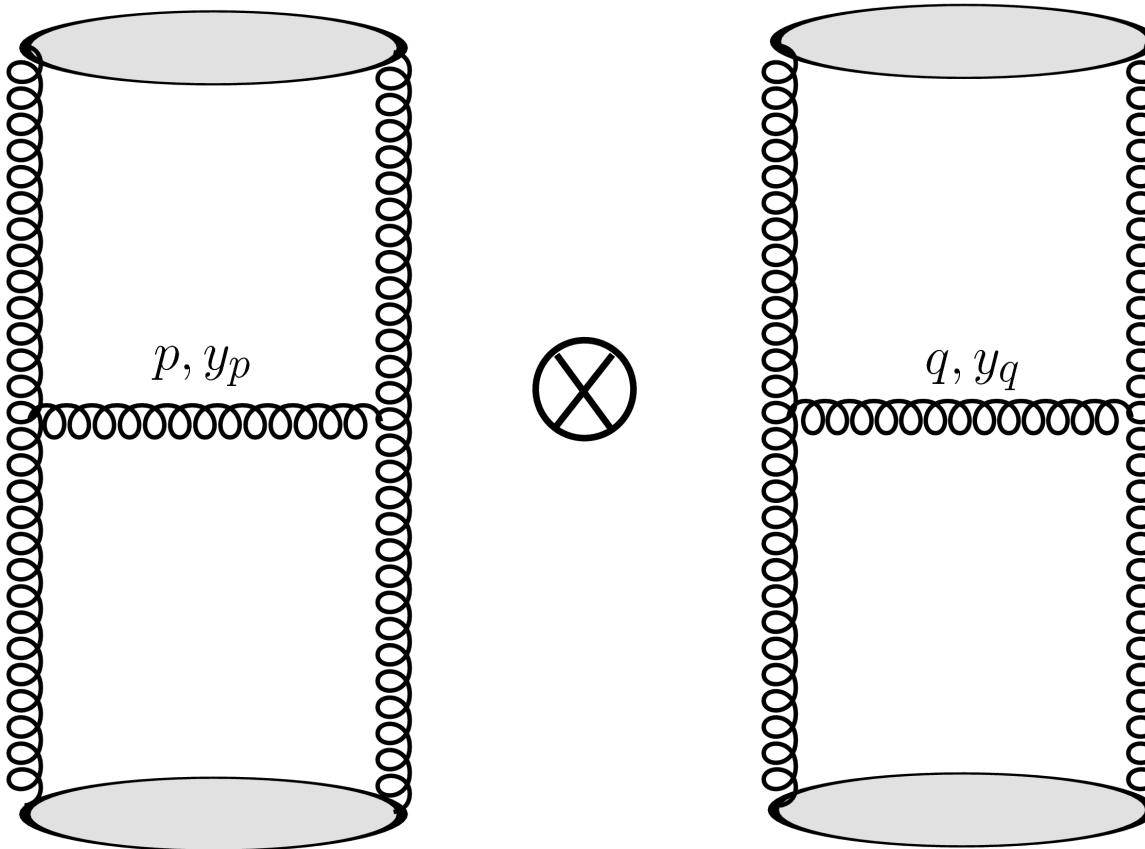
$$\left\langle \frac{dN_2}{dy_1 dy_2} \right\rangle_{\text{conn.}} = \frac{1}{k} \left\langle \frac{dN}{dy_1} \right\rangle \left\langle \frac{dN}{dy_2} \right\rangle$$



$k^{-1} x$



independent 2-gluon production with MV action

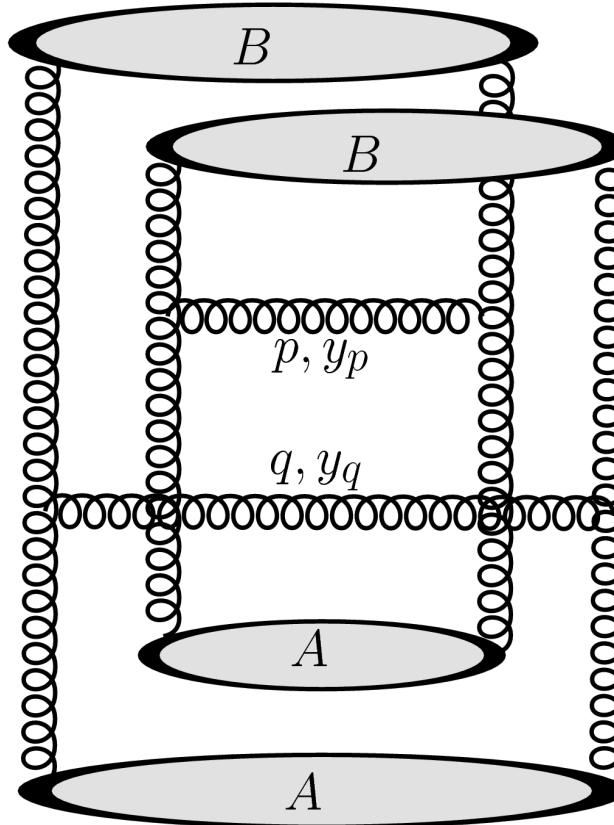


$$\sim N_c^2(N_c^2 - 1)^2 (\pi R^2)^2 \frac{(\mu^2)^4}{p^2 q^2} \int \frac{dk^2}{k^2} \frac{1}{(p - k)^2} \int \frac{dk'^2}{k'^2} \frac{1}{(q - k')^2}$$

$$\sim \frac{N_c^2(N_c^2 - 1)^2}{p^4 q^4} (\mu^2)^4 (\pi R^2)^2 \log \frac{p^2}{Q_s^2} \log \frac{q^2}{Q_s^2}$$

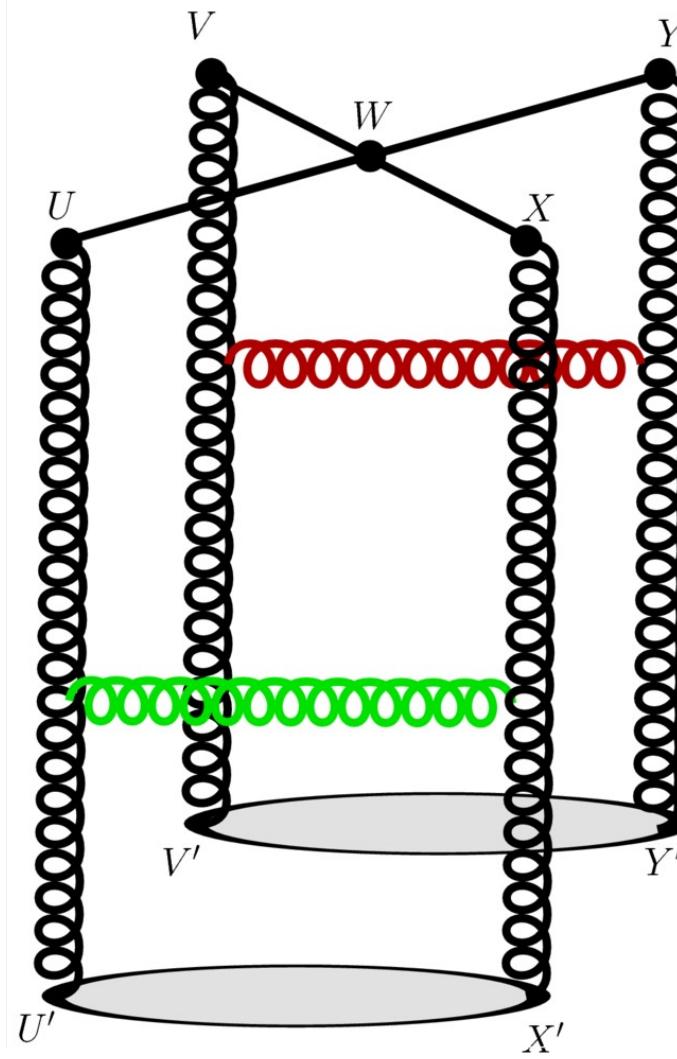
0

double gluon production with MV action



$$\sim N_c^2(N_c^2 - 1) \pi R^2 \frac{(\mu^2)^4}{p^2 q^2} \int \frac{dk^2}{k^4} \frac{1}{(p - k)^2} \frac{1}{(q - k)^2}$$

$$\sim \frac{N_c^2(N_c^2 - 1)}{p^4 q^4} (\mu^2)^4 \pi R^2 Q_s^{-2}$$



$$\sim -N_c^2(N_c^2 - 1)^2 \frac{(\mu^2)^2 (\mu^4)^2}{\kappa_4} \frac{\pi R^2}{p^2 q^2} \int \frac{dk^2}{k^2} \frac{1}{(p-k)^2} \int \frac{dk'^2}{k'^2} \frac{1}{(q-k')^2}$$

$$\sim -\frac{N_c^2(N_c^2 - 1)^2}{p^4 q^4} \frac{(\mu^2)^2 (\mu^4)^2}{\kappa_4} \pi R^2 \log \frac{p^2}{Q_s^2} \log \frac{q^2}{Q_s^2}$$

compute k^{-1} from 2-particle connected diagrams:

Dumitru + E.P. 2012

$$\frac{Q_s^2 S_\perp}{2\pi} \frac{1}{k} \simeq \frac{1}{N_c^2 - 1} - 3 \frac{N_c^2 + 1}{N_c^2 - 1} \beta + \dots$$

$$MV \quad \rho^4$$

$\beta > 0$ makes k bigger,
should be sufficiently small so as not to ruin KNO !

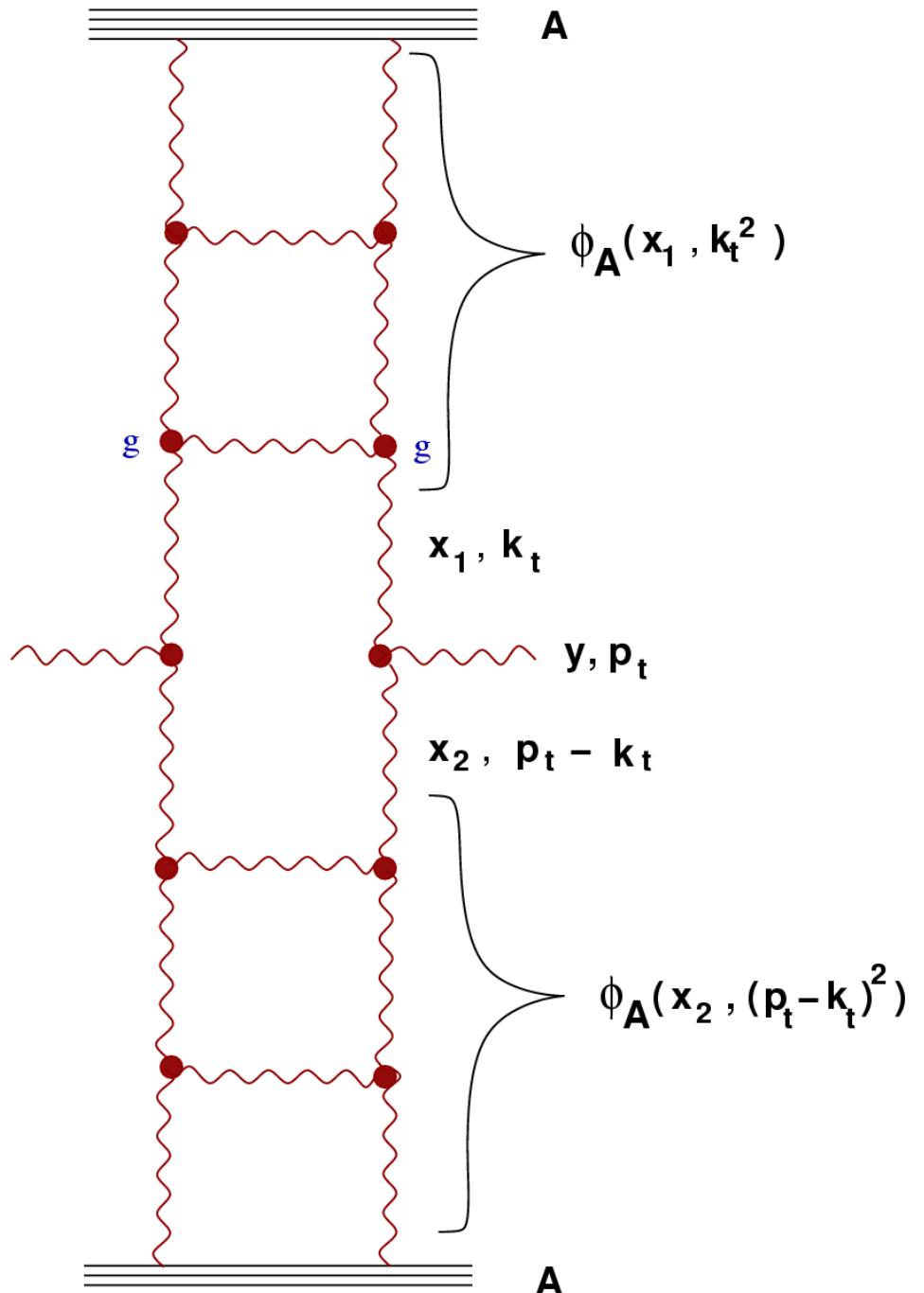
$$\beta \equiv \frac{C_F^2}{6\pi^3} \frac{g^8}{Q_s^2 \kappa_4} \left[\int_{-\infty}^{\infty} dz^- \mu^4(z^-) \right]^2$$

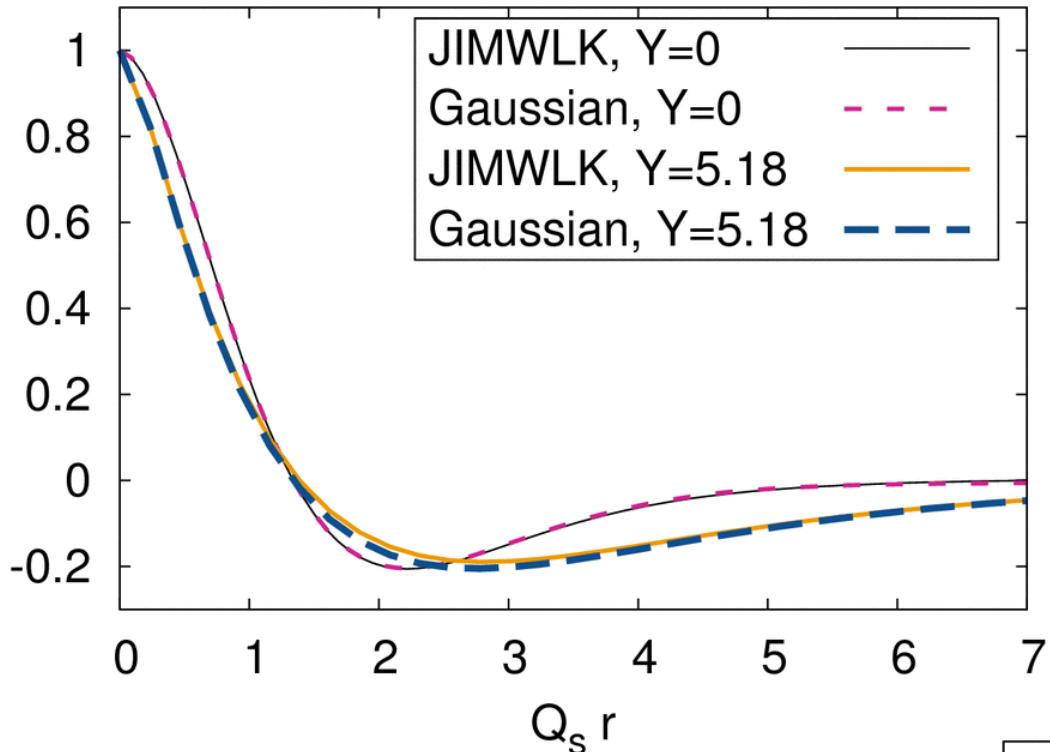
$$\approx 0.01 \text{ A}^{-2/3} \quad (\text{from fit to AAMQS dipole})$$

Evolution with energy

Fluctuations in evolution ladders:

do (rapidity-enhanced)
quantum fluctuations
satisfy KNO scaling ?

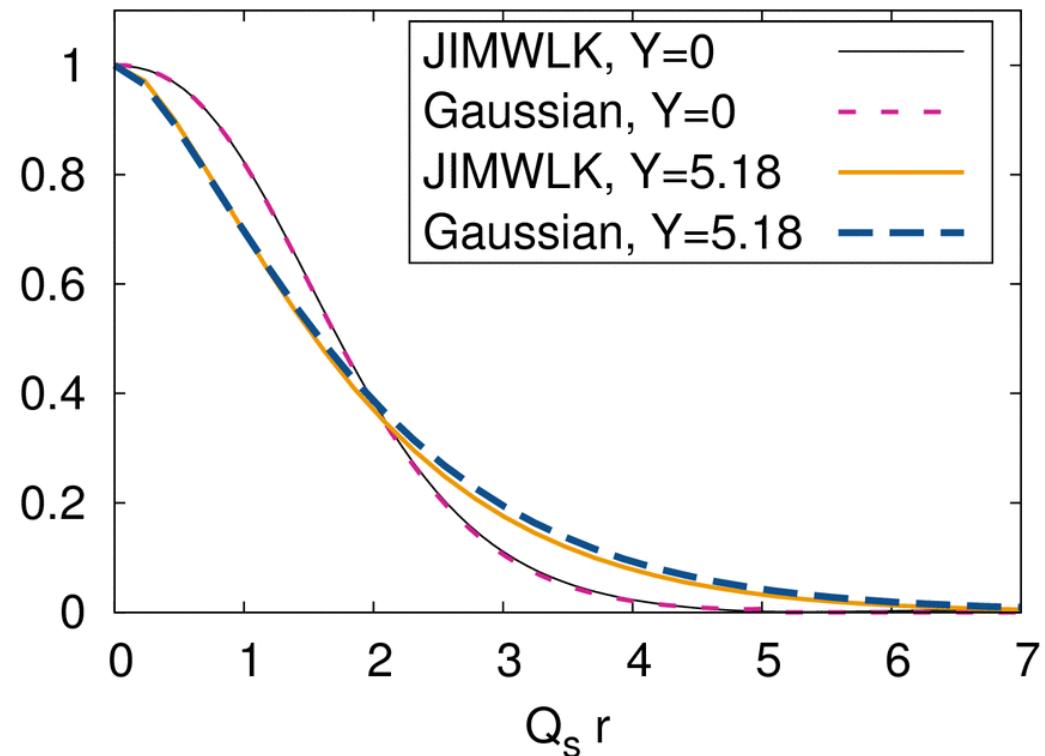


$\sigma_{\square}^{(r)}$ 

JIMWLK evolution appears to preserve Gaussianity of initial MV action ! (more to come)

Quadrupole evolution

$$Q \equiv \frac{1}{N_c} \langle \text{tr} V_x V_y^\dagger V_u V_w^\dagger \rangle$$

 $\sigma_{\square}^{(r)}$ 

*A.D., Jalilian-Marian, Lappi,
Schenke, Venugopalan:
arXiv:1108.4764*

Summary

- multiplicity distributions in pp @ LHC exhibit KNO scaling ($\eta=0$, $n/\bar{n} < \sim 3$)
- can be described by NBD with $k \ll \bar{n}$
- approx. KNO scaling predicted for p+Pb @ LHC (slight distortion of KNO due to Glauber flucs)
- higher-order eccentricities ε_3 etc. in HIC increase
- theoretical studies of fluctuations:
 - constrain magnitude of higher p^n operators
 - evolution with energy to test validity of Gaussian approximation from JIMWLK is in progress with B. Schenke, R. Venugopalan

Backup Slides

k_\perp -factorization, multiplicity in A+B $\rightarrow g+X$

(generalized) unintegrated gluon distribution:

$$\varphi(k, Y; b, A) = \frac{C_F k^2}{\alpha_s(k)} \int \frac{d^2 \mathbf{r}}{(2\pi)^3} e^{-i \mathbf{k} \cdot \mathbf{r}} \mathcal{N}_A(r, Y; b, A)$$

multiplicity: (Kharzeev, Levin, Nardi ansatz)

$$\frac{dN^{A+B \rightarrow g}}{dy d^2 b} = K \frac{1}{2C_F} \int \frac{d^2 p_t}{p_t^2} \int^{p_t} d^2 k_t \alpha_s(Q) \varphi \left(\frac{|p_t + k_t|}{2}, x_1 \right) \varphi \left(\frac{|p_t - k_t|}{2}, x_2 \right)$$

- finite at $p_t \rightarrow 0$ if UGD does not blow up
- $x_{1,2} = (p_t/\sqrt{s}) \exp(\pm y)$; $Y_{1,2} = \log(x_0/x_{1,2})$
where $x_0=0.01$ is assumed onset of rcBK evol.

rcBK evolution:

basic “degrees of freedom”: dipole scattering amplitude in fund. rep. (2-point fct)

$$\mathcal{N}_F(r, Y; b, A) \equiv \frac{1}{N_c} \text{tr} \langle 1 - V^\dagger(y) V(z) \rangle_Y$$

$$\mathbf{r} = \mathbf{y} - \mathbf{z}$$

BK equation (incl. non-linear terms → saturation of scattering amplitude!)

$$\frac{\partial \mathcal{N}(r, Y)}{\partial Y} = \int d^2 r_1 K(r, r_1, r_2) [\mathcal{N}(r_1, Y) + \mathcal{N}(r_2, Y) - \mathcal{N}(r, Y) - \mathcal{N}(r_1, Y) \mathcal{N}(r_2, Y)]$$

$$\mathbf{r} = \mathbf{r}_1 - \mathbf{r}_2$$

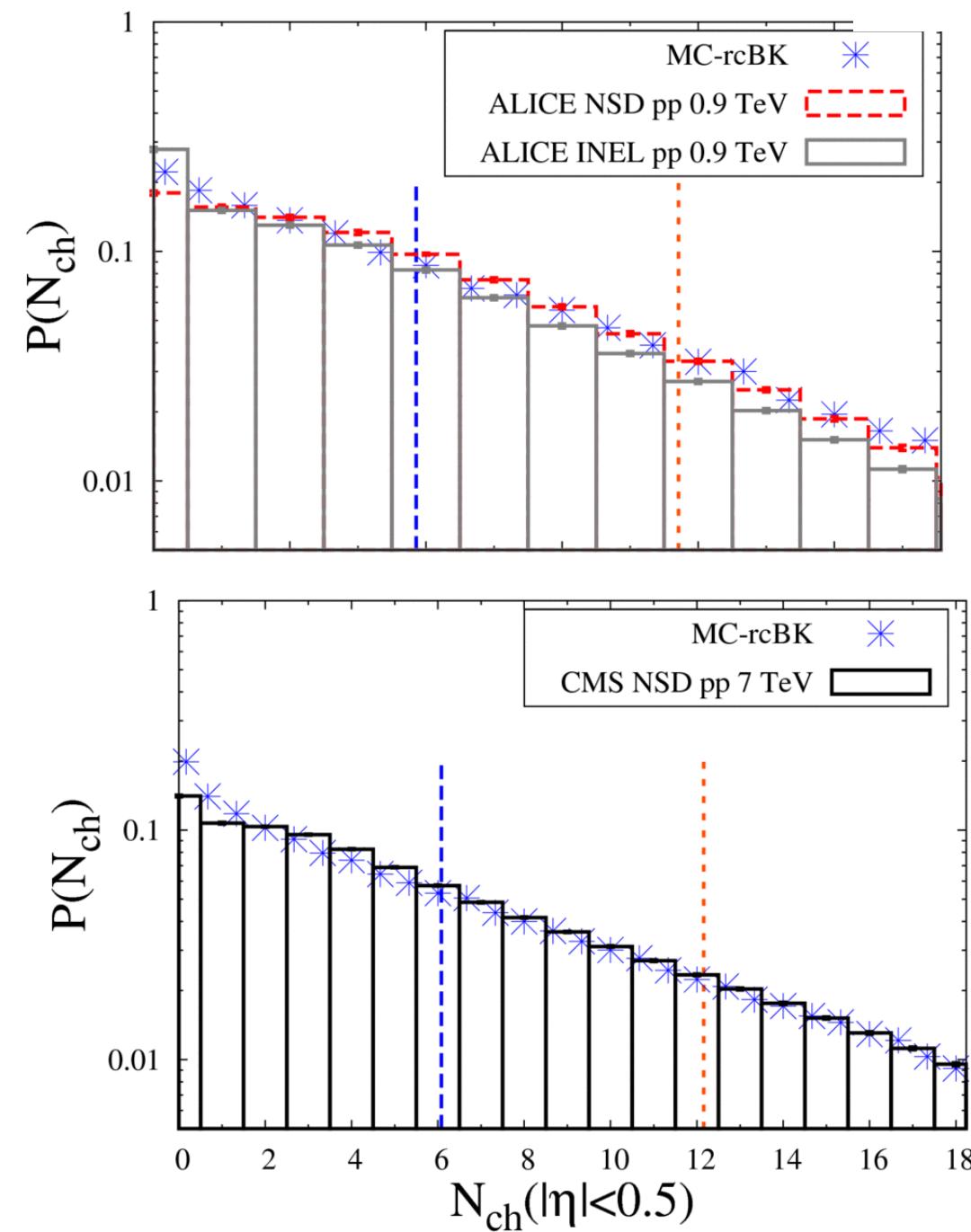
running-coupling kernel (Balitsky prescription)

$$K(\mathbf{r}, \mathbf{r}_1, \mathbf{r}_2) = \frac{N_c \alpha_s(r^2)}{2\pi^2} \left[\frac{1}{r_1^2} \left(\frac{\alpha_s(r_1^2)}{\alpha_s(r_2^2)} - 1 \right) + \frac{r^2}{r_1^2 r_2^2} + \frac{1}{r_2^2} \left(\frac{\alpha_s(r_2^2)}{\alpha_s(r_1^2)} - 1 \right) \right]$$

dipole scattering amplitude in
adj. rep.

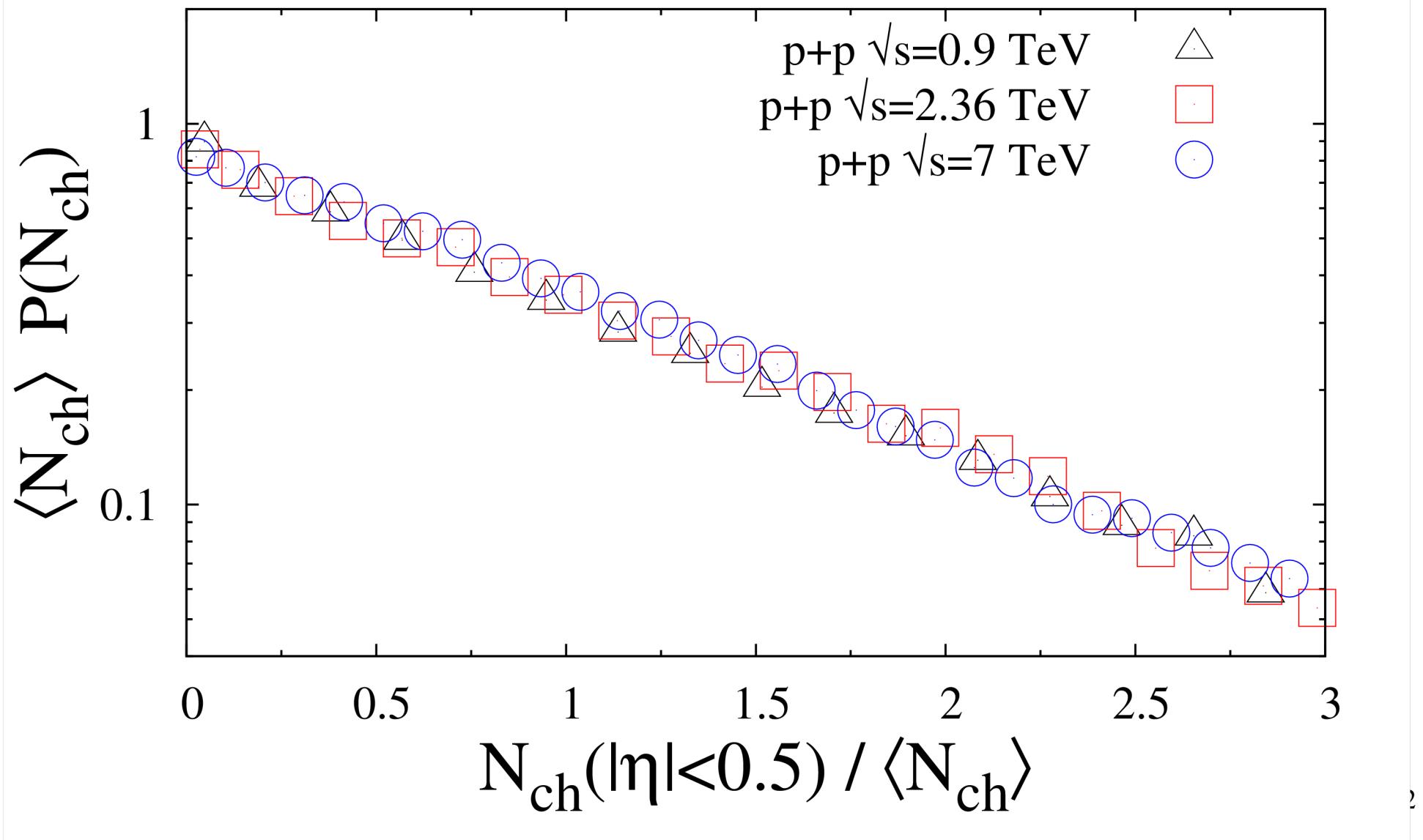
$$\mathcal{N}_A = 2 \mathcal{N}_F - \mathcal{N}_F^2$$

result for constant $k = \frac{1}{\pi} \Delta x_{\perp}^2 \Delta \eta \Lambda_{\text{QCD}}^2$



energy dependent $k \sim E_{\text{CM}}^{0.2}$

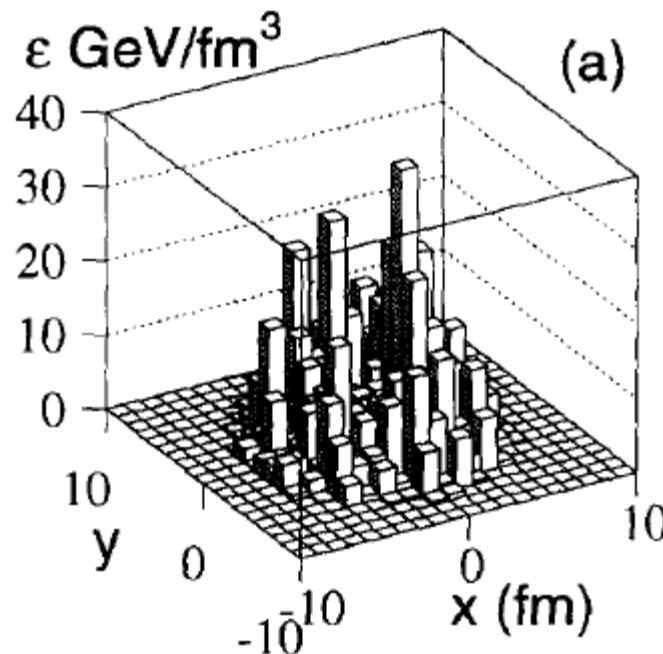
MC-rcBK, KNO scaling with $k \propto (\sqrt{s} / 900\text{GeV})^{0.2}$



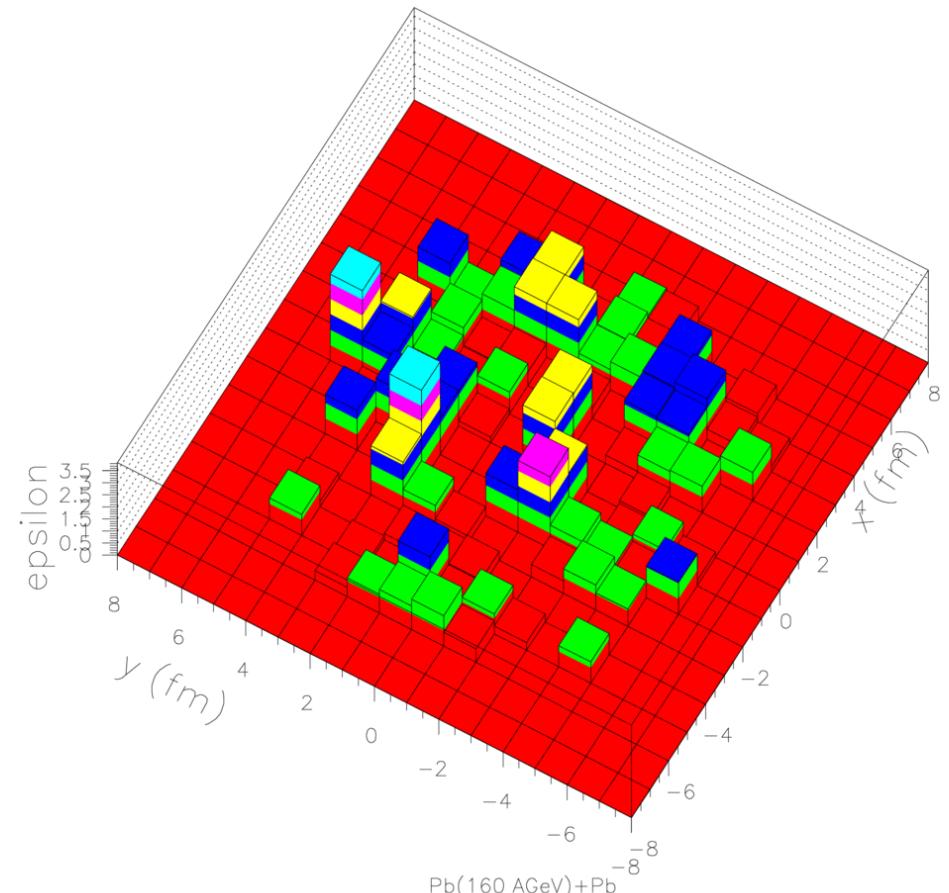
Non-Gaussian initial conditions for high-energy evolution

- Odderon operator $-d^{abc}\rho^a\rho^b\rho^c/\kappa_3$ S. Jeon and R. Venugopalan,
Phys. Rev. D70, 105012 (2004); 71, 125003 (2005)
- Effective action for a system of $k \sim N_c A^{1/3} \gg 1$ valence quarks in SU(3);
- Random walk of SU(3) color charges in the space of representations (m,n);
- Probability $P(m, n) = e^{-S(m, n)}$
$$S(m, n; k) \simeq \frac{N_c}{k} C_2(m, n) - \frac{1}{3} \left(\frac{N_c}{k} \right)^2 C_3(m, n) + \frac{1}{6} \left(\frac{N_c}{k} \right)^3 C_4(m, n)$$
- Casimir operator C_2, C_3, C_4 ntation (m,n)
- Define color charge per unit area $\rho^a \equiv g Q^a / \Delta^2 x$
where $|Q| = \sqrt{Q^a Q^a} \equiv \sqrt{C_2}$

Lumpy initial conditions in AA collisions

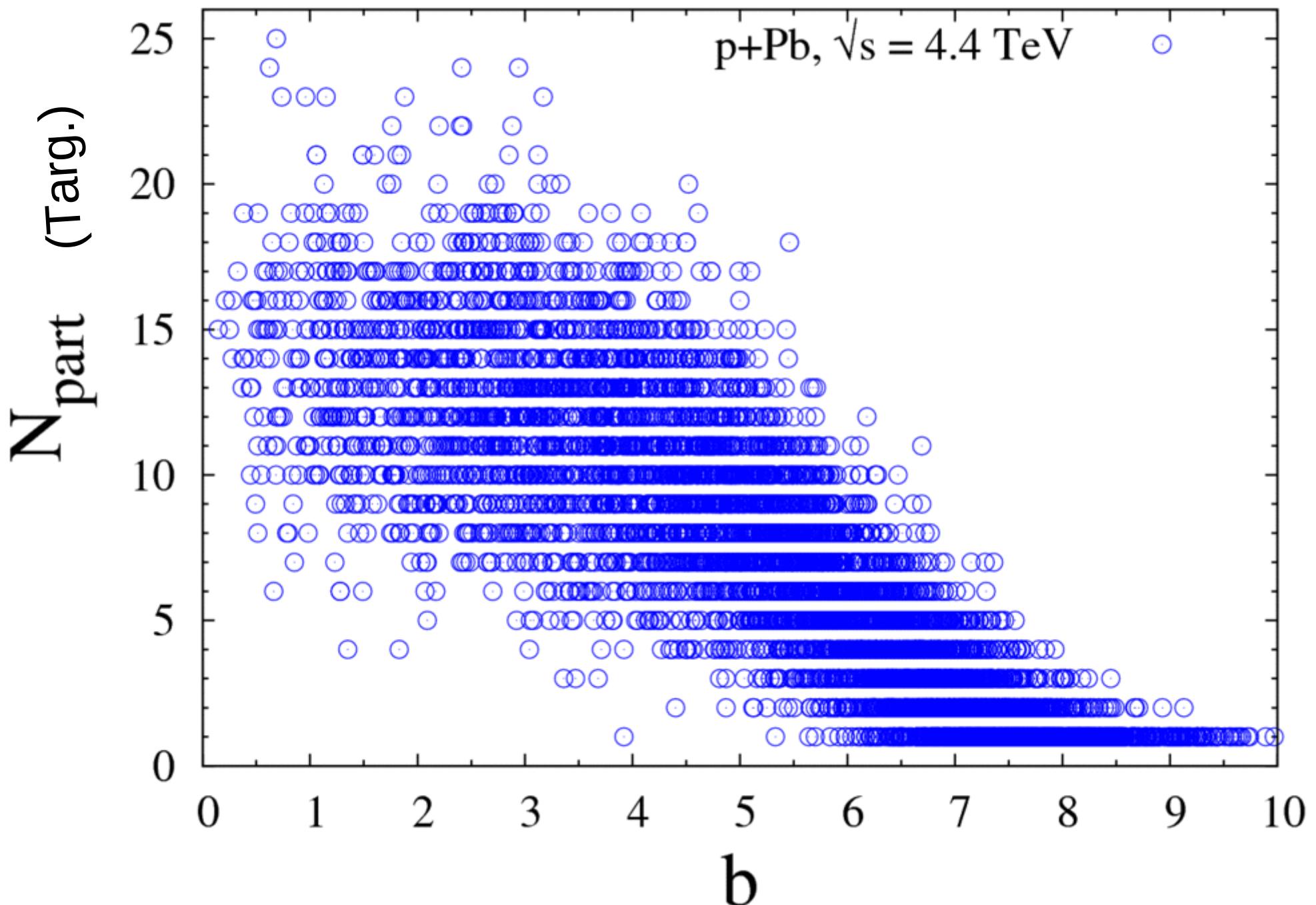


Gyulassy, Rischke, Zhang:
NPA 613 (1997)



M. Bleicher *et al.*: QM97 (Tsukuba)
NPA 638 (1998) p.391

N_{part} fluctuations in p+Pb:



$$\psi(z;\beta) / \psi(z;\beta=0)$$

